Symmetry algebras for superintegrable systems

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Abstract

It is shown that the symmetry algebra of quantum superintegrable system can be always chosen to be u(N), N being the number of degrees of freedom.

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Classical dynamical system of N degrees of freedom is called maximally superintegrable if it admits 2N-1 independent, globally defined integrals of motion. They form the complete set of integrals and any other integral of motion can be expressed in terms of them. In particular, the Poisson bracket of two basic integrals, being again a constant of motion, is (in general, nonlinear) their function. We conclude that the integrals of motion for maximally superintegrable system form a finite W-algebra [1], [2].

An interesting question arises whether this algebra can be linearized to Lie algebra by a judicious choice of basic integrals. The answer to this question, for a large class of confining systems, admitting action-angle variables is positive [3] (see Ref. [4] for older papers on the subject, both in classical and quantum theory). Moreover, assuming that <u>all</u> action variables do appear in the hamiltonian, the Lie algebra obtained is universally u(N). The price one has to pay for this result is that the resulting integrals are usually quite complicated functions of basic dynamical variables.

The aim of the present note is to extend this result to the quantum case.

The first problem which arises when passing to the quantum case is the very definition of quantum superintegrable system. We shall use the following one which seems to cover most intresting cases. Assume we have a set of commuting selfadjoint operators \hat{I}_k , $k=1,\ldots,N$ (quantum actions) such that:

- (i) the spectrum of each \hat{I}_k consists of the eigenvalues of the form $n_k + \sigma_k$, where $n_k \in N$ is any natural number and $\sigma_k \in \mathbf{R}$ is fixed.
- (ii) the common eigenvectors of all \hat{I}_k , | n >, $n \equiv (n_1, \ldots, n_N)$, span the whole space of states
- (iii) the hamiltonian can be written as

$$\hat{H} = H(\sum_{k=1}^{N} m_k \hat{I}_k) \tag{1}$$

with $m_k \in N$, $k = 1, \ldots, N$; we shall also assume that all $m_k \neq 0$. Let us also note that we can assume that m_1, \ldots, m_N have no common divisor except one.

We believe the above definition is flexible enough to cover known cases. Sometimes it may appear necessary to take an orthogonal sum of few Hilbert spaces described above to cover the whole space of states (the simplest example semms to be the planar Kepler problem).

The vectors $| n \rangle$ are eigenvectors of \hat{H} ,

$$\hat{H} \mid n \rangle = E_n \mid n \rangle$$

$$E_n = H \left(\sum_{k=1}^N m_k n_k + \sum_{k=1}^N m_k \sigma_k \right)$$
(2)

We see that the energy spectrum is degenerate: all eigenvectors $|n\rangle$ with $\sum_{k=1}^{N} m_k n_k$ fixed give the same energy. To classify the degeneracy according to the representations of some symmetry algebra we make the following trick [5]. Let m be the least common multiple of all m_k , $k = 1, \ldots, N$, and let $l_k = \frac{m}{m_k}$. Moreover, let us put

$$n_k = q_k l_k + r_k, \ 0 \le r_k \le l_k - 1, \ k = 1, \dots, \ N$$
 (3)

Then the expression (2) for energy eigenvalues takes the form

$$E_n = H\left(m\sum_{k=1}^{N} q_k + \sum_{k=1}^{N} m_k (r_k + \sigma_k)\right)$$
 (4)

Let $r \equiv (r_1, \ldots, r_N)$ be fixed; define X_r to be the subspace spanned by all vectors $| n \rangle$ such that $n_k = q_k l_k + r_k$; within X_r one can make an identification: $| n \rangle \equiv | q \rangle$, $q \equiv (q_1, \ldots, q_N)$. The whole space of states is then the orthogonal sum of $l_1 \cdot l_2 \cdot \ldots \cdot l_N$ subspaces X_r ,

$$X = \bigoplus_r X_r \tag{5}$$

When restricted to X_r energy spectrum of \hat{H} takes particularly simple form described by eq. (4). The relevant symmetry algebra can be constructed as follows. First, we define the creation and annihilation operators a_i^+ , a_i , $i = 1, \ldots, N$, as follows

$$a_i^+ \mid (n_1, \ldots, n_i, \ldots, n_N) > = \sqrt{n_i + 1} \mid (n_1, \ldots, n_i + 1, \ldots, n_N) >$$

 $a_i \mid (n_1, \ldots, n_i, \ldots, n_N) > = \sqrt{n_i} \mid (n_1, \ldots, n_i - 1, \ldots, n_N) >$ (6)

On the other hand, in each X_r one can define another set of creation and annihilation operators

$$b_{ri}^{+} \mid (q_1, \ldots, q_i, \ldots, q_N) > = \sqrt{q_i + 1} \mid (q_1, \ldots, q_i + 1, \ldots, q_N) > b_{ri} \mid (q_1, \ldots, q_i, \ldots, q_N) > = \sqrt{q_i} \mid (q_1, \ldots, q_i - 1, \ldots, q_N) >$$
 (7)

Obviously, the operators b_{ri}^+ , b_{ri} are expressible in terms of basic ones a_i^+ , a_i ; the relevant formulae read:

$$b_{ri} = \left(\frac{N_i + l_i - r_i}{l_i}\right)^{\frac{1}{2}} \Pi_{s=1}^{l_i} (N_i + s)^{-\frac{1}{2}} \cdot a_i^{l_i}$$

$$b_{ri}^+ = \left(\frac{N_i - r_i}{l_i}\right)^{\frac{1}{2}} \Pi_{s=1}^{l_i} (N_i - s + 1)^{-\frac{1}{2}} \cdot (a_i^+)^{l_i}$$
(8)

where $N_i \equiv a_i^+ a_i$ are the standard particle-number operators.

Now, the operators b_{ri} , b_{ri}^+ can be used in order to build symmetry algebra in each X_r . To this end let λ_{α} , $\alpha = 1, \ldots, N^2$, form the basis of u(N) algebra. Then the operators

$$\hat{\Lambda}_{r\alpha} \equiv \sum_{i, j=1}^{N} b_{ri}^{+}(\lambda_{\alpha})_{ij} b_{rj} \tag{9}$$

have the following properties: (a) they are hermitean, (b) commute with \hat{H} in X_r , (c) obey u(N) commutation rules in X_r .

Having constructed u(N) symmetry algebra in each X_r one can define the symmetry operators in the whole space of states X. Let Q_r be the orthogonal projector on X_r . Then the operators

$$\hat{\Lambda}_{\alpha} = \sum_{r} Q_{r} \hat{\Lambda}_{r\alpha} Q_{r} \equiv \sum_{r} \hat{\Lambda}_{r\alpha} Q_{r} \tag{10}$$

span the u(N) symmetry algebra in the total space of states X.

It remains to construct the projectors Q_r . It is easy to check that the following operators do the job

$$Q_r \equiv \prod_{i=1}^{N} P_i$$

$$P_i = \frac{1}{l_i} \sum_{s=0}^{l_i - 1} e^{\frac{2i\pi s}{l_i} (N_i - r_i)}$$
(11)

The creation and anihilation operators used above were constructed in a quite general abstract way. However, an interesting question is whether they can be expressed in terms of basic dynamical variables. In most cases the answer to this question is positive although the resulting expressions are, in general, very complicated. Indeed, in almost all interesting cases (like, for example, interacting particles, with or without spin, described by the natural hamiltonian) the space of states carries an irreducible representation of the algebra of observables. The only possible exceptions we can see are the existence of superselection rules or the cooexistence of discrete and continuum spectrum of the hamiltonian. In the latter case we must restrict ourselves to the discrete part of spectrum. The basic dynamical variables do have, in general, nonvanishing matrix elements between eigenstates coresponding to discrete and continuous spectra. This can be cured, for example, by inserting the projector on discrete subspace but this makes the whole procedure even more complicated.

As we have mentioned above we are convinced that our assumptions (i)-(iii) are flexible enough to cover known cases of superintegrable systems. Below we give a number of examples.

We start with the most obvious one - the harmonic oscillator with rational frequency ratios. The relevant hamiltonian reads

$$H = \sum_{k=1}^{N} \left(\frac{p_k^2}{2m_k} + \frac{m_k^2 l_k^2 \omega^2}{2} x_k^2\right) = \sum_{k=1}^{N} H_k, \tag{12}$$

Defining

$$\hat{I}_k = \frac{1}{\hbar\omega l_k} H_k \tag{13}$$

we easily see that (i) - (iii) hold with all $\sigma_k = \frac{1}{2}$.

We can generalize this by adding the inverse square term [6]

$$H = \sum_{k=1}^{N} \left(\frac{p_k^2}{2m_k} + \frac{m_k^2 l_k^2 \omega^2}{2} x_k^2 + \frac{\alpha_k^2}{x_k^2}\right) = \sum_{k=1}^{N} H_k; \tag{14}$$

then

$$\hat{I}_k = \frac{1}{2\hbar\omega l_k} H_k \tag{15}$$

and $\sigma_k = \frac{3}{4}$ for all k do the job.

The next example is the twodimensional Kepler problem

$$H = \frac{\vec{p}^2}{2\mu} - \frac{\alpha}{r} \tag{16}$$

The spectrum is descrete for negative energies and reads

$$E_n = \frac{-\alpha \mu^2}{2\hbar^2 (n+|m|+1)^2},\tag{17}$$

m being the value of angular momentum. Therefore, we restrict ourselves to the subspace spanned by negative-energy eigenvectors of (16). Moreover, due to the appearance of |m| in eq.(17) we split this subspace futher accordingly to whether m is nonnegative or negative.

The dynamical symmetry behind is obtained by defining Runge-Lenz vector [7]

$$A_i = \frac{1}{2\mu} \varepsilon_{ik} (p_k L + L p_k) - \frac{\alpha}{r} x_i. \tag{18}$$

One finds the following algebra

$$[L, A_i] = i\hbar \varepsilon_{ik} A_k$$

$$[A_i, A_j] = -\frac{2i\hbar}{\mu} \varepsilon_{ij} HL$$
(19)

This is a quadratic algebra. However, in the subspace of constant negative energy it becomes sO(3) algebra.

In order to see that (i) - (iii) are fulfiled consider the subspace corresponding to $m \geq 0$ (the case m < 0 can be treated along the same lines). Define

$$\hat{I}_1 = \frac{1}{\hbar}L$$

$$\hat{I}_2 = \frac{1}{\hbar}(\mu\sqrt{\frac{-\alpha}{2H}} - L)$$
(20)

These operators are well-defined, selfadjoint and commuting; their spectra obey (i) with $\sigma_1 = 0$, $\sigma_2 = \frac{1}{2}$. Moreover,

$$H = \frac{-\alpha\mu^2}{2\hbar^2(\hat{I}_1 + \hat{I}_2)^2} \tag{21}$$

so that $m_1 = m_2 = 1$ in eq.(1).

The three dimensional Kepler problem can be deal with in the same way. The orbital angular momentum \vec{L} , together with the Runge - Lenz vector [7]

$$\vec{A} = \frac{1}{2\mu} (\vec{p} \times \vec{L} - \vec{L} \times \vec{P}) - \frac{\alpha}{r} \vec{r}$$
 (22)

form the quadratic algebra

$$[L_{i}, L_{j}] = i\hbar \varepsilon_{ijk} L_{k}$$

$$[L_{i}, A_{j}] = i\hbar \varepsilon_{ijk} A_{k}$$

$$[A_{i}, A_{j}] = -\frac{2ihbar}{\mu} \varepsilon ijk L_{k} H$$
(23)

Again, let us consider the negative energy subspace. Define

$$\hat{I}_{1} = \sqrt{\frac{\vec{L}^{2}}{\hbar^{2}} + \frac{1}{4}} + \frac{L_{3}}{\hbar}$$

$$\hat{I}_{2} = \sqrt{\frac{\vec{L}^{2}}{\hbar^{2}} + \frac{1}{4}} - \frac{L_{3}}{\hbar}$$

$$\hat{I}_{3} = \frac{\mu}{\hbar} \sqrt{-\frac{\alpha}{2H}} - \frac{1}{2} (\hat{I}_{1} + \hat{I}_{2})$$
(24)

Then (i) - (iii) are obeyed with $\sigma_1 = \sigma_2 = \sigma_3 = \frac{1}{2}$ and

$$H = -\frac{2\alpha\mu^2}{\hbar^2(\hat{I}_1 + \hat{I}_2 + 2\hat{I}_3)^2}$$
 (25)

Both harmonic oscillator and Kepler problem admit generalization consisting in replacing the Euclidean configuration space by the sphere [8] which preserves superintegrability. It is easy to check using the results of [8] that the relevant operators \hat{I}_k can be constructed in a similar way as in the "plane" case. Therefore, both models fit into our scheme.

Let us note in passing that our method applies to both models in arbitrary number of dimensions. This is because the dynamical symmetries are here SU(n) or SO(n+1) and one can use the Gelfand-Tseytlin method to construct the operators \hat{I}_k .

As a final example let us take the rational Calogero model [9]. It is known to be superintegrable also on the quantum level [10]. Using the results of [11] it is not difficult to check that our assumptions (i) - (iii) can be satisfied by an appropriate choice of \hat{I}_k operators.

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